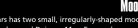
# Express mission to Mars

Mars Express is the European Space Agency's first mission to Mars. Its role is to survey the planet from the topmost layer of the atmosphere down to beneath the surface. Since its arrival at Mars, the mission has proved a great success and has given us many new insights into the Martian environment.



Phobos and Deimos. Mars Express has provided the most complete detailed picture of Phobos to date. The OMEGA instrument has mapped the chemical tion of the moons' surfaces and minerals that have been exposed at the bottom of impact craters



Highlights include the detection of methane in the Martian atmosphere, the discovery of water ice reserves buried under the planet's surface and the discovery that aurora occur in the upper Martian atmosphere.

The mission consisted of the Mars Express orbiter and the

Mars Express, which will operate until at least May 2009, orbits Mars in 6 hours and 43 minutes and follows an elliptical path that gives it close up and global views of the planet.



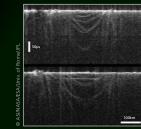
# Versatile vision the many eyes of Mars Express

The Mars Express orbiter carries seven instruments that are tailor-made to monitor different aspects of the Martian environment. HRSC, OMEGA and MARSIS study the planet's surface or subsurface, and PFS, SPICAM and ASPERA monitor the Martian atmosphere and its interaction with the solar wind. The Mars Radio Signal experiment uses radio transmissions between the spacecraft and the Earth to deduce information about the planet's surface and interior. UK scientists are involved in the ASPERA, MARSIS and HRSC instruments.

HRSC colour image of the caldera of Olympus Mons

HRSC black & white image of the eastern scarp of Olympus Mons

PFS and SPICAM instruments analyse infrared and ultraviolet radiation to determine the composition of



Advanced Radar for Subsurface and lonosphere Sounding) probes up to 5 kilometres oouncing radio signals off layers in the rock and analysing above appear to show a buried impact crater filled with a



◆ ASPERA-3 consists of 4 sensors that the Martian atmosphere with the solar wind

OMEGA analyses the light reflected from the surface of Mars to create a map of rock-types on the surface and areas covered by frost or ice. Different chemicals absorb light strongly at characteristic wavelengths. These three images of the Martian south pole show (from left to right) areas covered by carbon dioxide ice,

## Further Information

the Mars Express mission, try the following resources:

#### **World Wide Web sites**

www.esa.int/SPECIALS/Mars Express/

## www.uk2planets.org.uk

Patrick Moore on Mars Cassel, 2006

Roving Mars: Spirit, Opportunity and the Exploration of the Red Planet

The Real Mars

Joseph M Boyce
The Smithsonian Book of Mars

Heather Couper & Nigel Henbest Mars Headline, 2001

The Cambridge Photographic Guide to the Planets

Cover image: Credit: Viking Project, USGS, NASA



### Science & Technology Facilities Council

The Science and Technology Facilities Council operates world-class, large-scale research facilities; supports scientists and engineers world-wide; funds researchers in universities and provides strategic

The Council's Science in Society unit offers a wide range of support for teachers, scientists and communicators to facilitate greater engagemen with STFC science which includes astronomy science, particle physics and nuclear physics

Free Publications suitable for students aged 10-18.

- A Moon rock and meteorite loan scheme
- Oxfordshire and Edinburgh plus CERN in Geneva.
- partnership with young scientists who have been trained to support the teacher, act as role models and introduce their research.

### For scientists

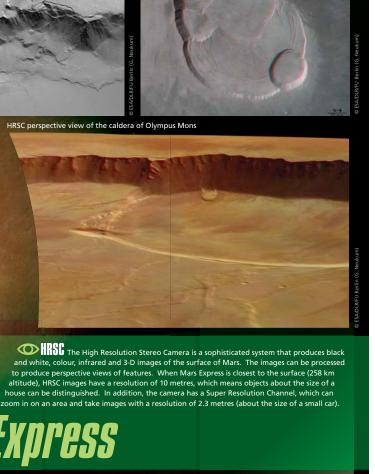
funding schemes and Fellowships for public

For further information telephone 01793 442175 or email gareth.james@stfc.ac.uk









HRSC 3-D image of the caldera of Olympus Mons